
**ANALOGIES, NEW PARADIGMS, AND OBSERVATIONAL
DATA AS GROWING FACTORS OF RELATIVISTIC
ASTROPHYSICS**

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Patterns in the scientific development of relativistic astrophysics are analyzed with special attention to the physics and astrophysics of black holes and gamma-ray bursts.

1 Introduction

Hagen Kleinert has been a pioneer in establishing analogies among widely separated fields of theoretical physics, applying relativistic quantum field theoretical techniques, notoriously his classical work with Richard Feynman [1], to the treatment of a variety of research topics ranging from condensed matter to crystal melting, polymer physics, phase transitions, differential geometry etc. I have always been very impressed by his profound knowledge of physics and his courage in approaching so vast a research program, as testified also in his classical books [2–4]. On the occasion of his sixtieth birthday, I am dedicating to him a few brief considerations based on my work, of how the role of analogies among different fields, the establishment of new paradigms, as well as the crucial and timely arrival of observational data have marked the development of relativistic astrophysics.

2 The Birth of Relativistic Astrophysics

In 1931, in analogy with the development made in the field of atomic physics by Enrico Fermi and his school [5], Lev D. Landau [6] was led to introduce

a new paradigm in the approach of astrophysical problems. This was not based on a hypothesis chosen merely for mathematical convenience, but was founded on new concepts developed in theoretical physics. Using the Fermi-Dirac statistics to study the equilibrium of a star paved the way to analyze the latest phases of the evolution of a star. This approach has developed into an entirely new theory for understanding white dwarfs in the classical book by Chandrasekhar [7]. Subsequently, new concepts were introduced by Julius Robert Oppenheimer, using the general relativistic techniques by Tolman and the seminal ideas of George Gamow (with whom Hagen Kleinert studied from 1965 to 1967 while working on his Ph.D. thesis). I recall here the treatment of neutron stars in the classic work with Oppenheimer's student Volkoff [8]. This entire research field reached full maturity with the discovery of pulsars in 1968, especially with the discovery of the crab nebula pulsar. The observation of the period of that pulsar and his rate of slowing-down gave clear evidence for identifying the first neutron star in the galaxy. It also confirmed our understanding that the energy source of pulsars was simply the rotational energy of the neutron star. The year 1968 can definitely be considered as the date of birth of relativistic astrophysics.

I was in Princeton in those days, initially at the university as a postdoctoral fellow in the group of John Archibald Wheeler, later as a member of the Institute for Advanced Study, and finally as an instructor and assistant professor at the university. The excitement about the neutron star discovery boldly led us directly to a yet unexplored classic paper by Oppenheimer and Snyder: *On Continued Gravitational Contraction* [9]. This opened up an entire new research field to which I have dedicated all the rest of my life and which is giving, still today, some distinctively important results. I will comment in the following on a few crucial moments, the way I remember them, that influenced very much the development of relativistic astrophysics, with particular emphasis on the establishment of analogies, new paradigms and crucial observational data.

3 Analogies Between Trajectories of Cosmic Rays and Trajectories in General Relativity

An “effective potential” technique had been used very successfully by Carl Størmer in the 1930s for studying the trajectories of cosmic rays in the earth's magnetic field [10]. In the fall of 1967, Brandon Carter visited Princeton and presented his remarkable mathematical work leading to the separability

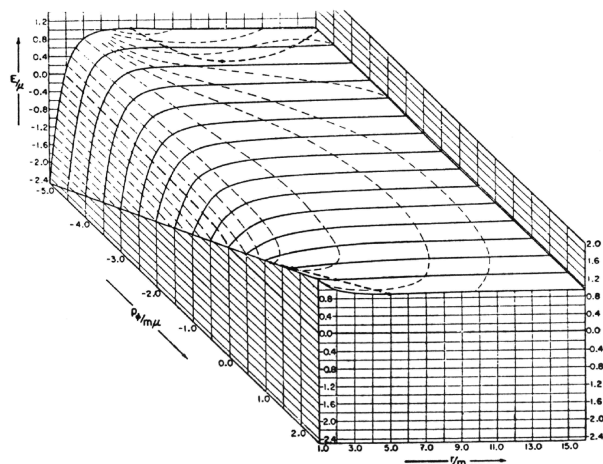


Figure 1. “Effective potential” around a Kerr black hole (reproduced from Ref. [13]).

of the Hamilton-Jacobi equations for the trajectories of charged particles in the field of a Kerr-Newmann geometry [11]. Carter's visit had a profound impact on our small group working with John Wheeler on the physics of gravitational collapse. Indeed, it was Johnny who had the idea to exploit the analogy between the trajectories of cosmic rays and the trajectories in general relativity, using Størmer's “effective potential” technique in order to obtain physical consequences from the set of first-order differential equations obtained by Carter. I still remember my preparing the $2m \times 2m$ grid plot of the effective potential for particles around a Kerr metric which finally appeared in print [12] (see Fig. 1). From this work came the celebrated result of the maximum binding energy of $1 - 1/\sqrt{3} \sim 42\%$ for corotating orbits, and $1 - 5/3\sqrt{3} \sim 3.78\%$ for counterrotating orbits in the Kerr geometry. We were very pleased to be associated with Brandon Carter in a “gold medal” award for this work presented by Yevgeny Lifshitz: in the last edition of Vol. 2 of the Landau and Lifshitz series, entitled *The Classical Theory of Fields*, both Brandon's and my work with Wheeler are mentioned as exercises for bright students! This was certainly a simple and fruitful analogy which led to a successful theoretical accomplishment, opening up a new window on completely unexpected general relativistic effects. It is interesting that it has become clear in recent years that the difference in the binding energies

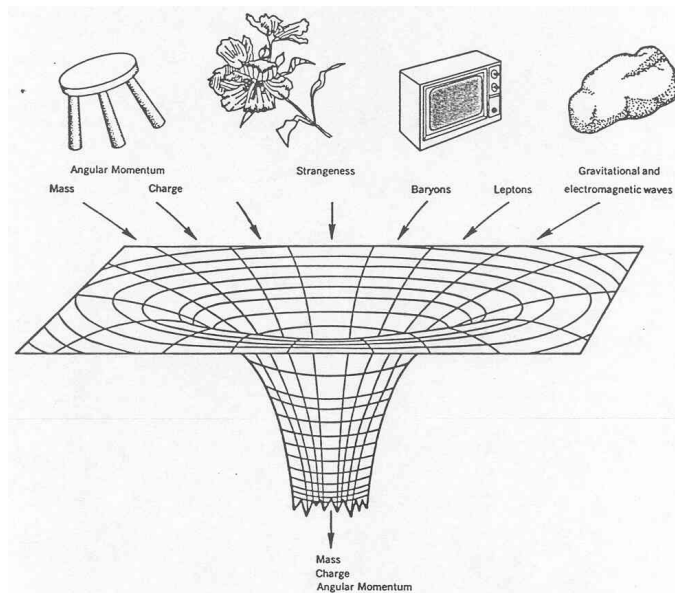


Figure 2. The black hole uniqueness theorem.

of the corotating and counterrotating orbits in a Kerr-Newmann geometry have become the object of direct astrophysical observations in binary X-ray sources.

4 Analogy of “Black Hole” with Elementary Physical System

In my 1971 article in *Physics Today* with Wheeler, entitled *Introducing the Black Hole*, we first proposed the famous “uniqueness theorem” stating that black holes can be completely characterized by their mass-energy E , charge Q , and angular momentum L [13]. This analogy between a black hole and a most elementary physical system was magnificently represented by Johnny in an unconventional figure in which TV sets, bread, flowers, and other objects loose their characteristic features and merge in the process of gravitational collapse into the three fundamental parameters of a black hole (see Fig. 2). That picture became the object of a great deal of light-hearted discussion in the physics community. A proof of this uniqueness theorem, satisfactory for some astrophysical cases, has been obtained after twenty five years of metic-

ulous mathematical work (see e.g. Regge and Wheeler [14], Zerilli [15,16], Teukolsky [17], Lee [18], and Chandrasekhar [19]). However, the proof still presents some outstanding difficulties in its most general form. Possibly, some progress will be reached in the near future with the help of computer algebraic manipulation techniques to overcome the extremely difficult mathematical calculations (see e.g. Cruciani [20], Cherubini and Ruffini [21], and Bini *et al.* [22,23]).

It is interesting that this proposed analogy, which appeared at first to be almost trivial, turned out to be one of the most difficult to be proven, implying a monumental work unsurpassed in difficulty, both in mathematical physics and relativistic field theories. An extremely good example of the difference between general relativity and classical physics is offered by the fact that this analogy is still unproven and that the most general perturbation of a black hole endowed with electromagnetic structure (EMBH) and rotation is still far from being solved. The solution of this problem from a mathematical physics point of view may have profound implications on our understanding of the fundamental physical laws.

5 Analogy Between Pulsars and Black Hole Physics: The Extraction of Rotational Energy

We were still under the sobering effects of the pulsar discovery and the very clear explanation by Tommy Gold and Arrigo Finzi that the rotational energy of the neutron star had to be the energy source of the pulsar phenomenon, when the first meeting of the European Physical Society took place in Florence in 1969. In a splendid talk, Roger Penrose [24] advanced the possibility that, much like in pulsars, the rotational energy could be in principle extracted in black holes. The first specific example of such an energy extraction process was given in a gedanken experiment using the above mentioned effective potential technique by Ruffini and Wheeler [25] (see Fig. 3), and later by Floyd and Penrose [26]. The reason for showing this figure here is a) to recall the first explicit computation and b) to remind the introduction of the “ergosphere” which is the region between the horizon of a Kerr-Newmann metric and the surface of an infinite redshift where the energy extraction process can occur, and finally c) to emphasize how contrived, difficult and conceptually novel such a mechanism of energy extraction can be. It is a phenomenon which is not localized at a point but can occur in an entire region: a global effect which relies essentially on the concept of a field. However, it can only

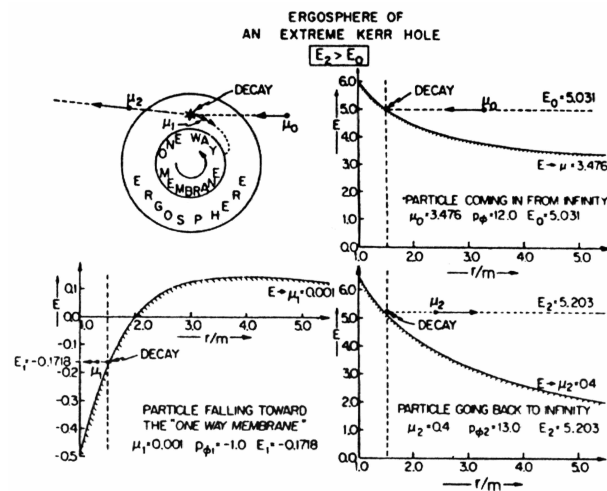


Figure 3. Decay of a particle of rest-plus-kinetic energy E_0 into a particle which is captured by the black hole with positive energy as judged locally, but negative energy E_1 as judged from infinity, together with a particle of rest-plus-kinetic energy $E_2 > E_0$ which escapes to infinity. The cross-hatched curves give the effective potential (gravitational plus centrifugal) defined by the energy E for constant values of p_ϕ and μ (reproduced from Ref. [27] with the kind permission of Ruffini and Wheeler).

work for very special parameters, and is in general associated with a reduction of the rest mass of the particle involved in the process. While it is almost trivial to slow down the rotation of a black hole and to increase its horizon by accretion of counterrotating particles, it is extremely difficult to extract the rotational energy from a black hole by a slowing-down process, as also clearly pointed out by the example in Fig. 3.

The establishment of this analogy offered us the opportunity to appreciate once more the profound difference of seemingly similar effects in general relativity and classical field theories. In addition we had the first glimpse to the existence of totally new phenomena, such as the dragging of the inertial frames around a rotating black hole, and thus to an entire new field of theoretical physics implied by the field equations of general relativity. The deep discussions of these problems with Demetrios Christodoulou, who was my first graduate student at the time in Princeton at the age of 17, led us to the discovery of the existence of “reversible and irreversible transformations” in black holes physics.

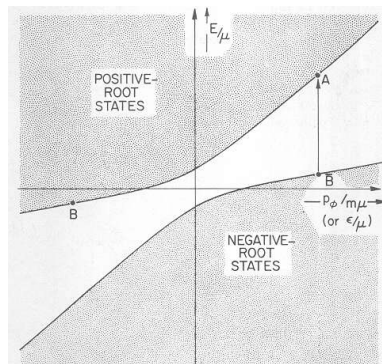


Figure 4. Reversing the effect of having added one particle A to the black hole by adding another particle B of the same rest mass but opposite angular momentum and charge in a “positive-root negative-energy state”. The addition of B is equivalent to the subtraction of B^- . Thus the combined effect of the capture of particles A and B is an increase in the mass of the black hole given by the vector B^-A . This vector vanishes and reversibility is achieved when and only when the separation between positive root states and negative root states is zero, in this case the hyperbolas coalesce to a straight line (reproduced from Ref. [28]).

6 The First Analogy Between Thermodynamics and Black Hole Physics: Reversible and Irreversible Transformations

It was by analyzing the capture of test particles by an electromagnetic black hole that we identified a set of limiting transformations which did not affect the surface area. These special transformations had to be performed very slowly, with a limiting value of zero kinetic energy on the horizon of the EMBH (see Fig. 4). It then became immediately clear that the total energy of an EMBH could in principle be expressed as a sum of “rest energy”, “Coulomb energy”, and “rotational energy”. The rest energy is “irreducible”, the other two are variable by the process of adding and extracting energy, respectively.

While Wheeler was mainly addressing the issue of the thermodynamical analogy, I and Demetrios were interested in the fundamental issue of the energetics of EMBH using the tools of the reversible and irreversible transformations. We finally obtained the general mass-energy formula for black holes [28]:

$$E^2 = M^2 c^4 = \left(M_{\text{ir}} c^2 + \frac{Q^2}{2\rho_+} \right)^2 + \frac{L^2 c^2}{\rho_+^2}, \quad (1)$$

$$S = 4\pi\rho_+^2 = 4\pi\left(r_+^2 + \frac{L^2}{c^2M^2}\right) = 16\pi\left(\frac{G^2}{c^4}\right)M_{\text{ir}}^2, \quad (2)$$

with the condition

$$\frac{1}{\rho_+^4}\left(\frac{G^2}{c^8}\right)(Q^4 + 4L^2c^2) \leq 1, \quad (3)$$

where M_{ir} is the irreducible mass, r_+ is the horizon radius, ρ_+ is the quasi-spheroidal cylindrical coordinate of the horizon evaluated at the equatorial plane, and S is the horizon surface area. Extreme black holes satisfy the equality in Eq. (3). The crucial point is that transformations at constant surface areas of the black hole, namely reversible transformations, can release an energy up to 29% of the mass-energy of an extremal rotating black hole and up to 50% of the mass-energy of an extremely magnetized and charged black hole. Since my Les Houches lectures *On the Energetics of Black Holes* [29], one of my main research goals has been to identify an astrophysical setting where the extractable mass-energy of the black hole could manifest itself. As we will see in the following, I propose that this extractable energy of an EMBH is the same as the energy source of gamma-ray bursts.

The thermodynamic analogy was further developed by Wheeler. By this time I had become convinced that the establishment of a one-to-one analogy between general relativistic effects and classical results was totally hopeless. It appeared to me that analogies could only be used as a very important tool to explore, possibly to be rationalized in a formula, the extremely vast theoretical world of the space-time structures contained in Einstein's theory of general relativity. A good example is our mass formula. Trying to enforce a perfect analogy is too risky and reductive.

7 Paradigm for Identifying the First “Black Hole” in Our Galaxy, and the Development of X-Ray Astronomy

The launch of the “Uhuru” satellite for the purpose of a first examination of the Universe in the X-ray spectrum, by a group directed by Riccardo Giacconi at the American Science and Engineering Department, marked a fundamental progress and generated a tremendous momentum in the field of relativistic astrophysics. The very fortunate work established simultaneous observations in the optical and in the radio wavelengths and thus allowed to have high-quality data on binary star systems composed of a normal star stripped off

matter by a compact massive companion star: either a neutron star or a black hole.

The *Maximum Mass of a Neutron Star* was the subject of the thesis of Clifford Rhoades, my second graduate student at Princeton. A criteria was found there to overcome fundamental ignorance about the behaviour of matter at supranuclear densities by establishing an absolute upper limit to the neutron star mass. Our result was only based on general relativity, causality, and the behaviour of matter at nuclear and subnuclear densities. This work, presented at the 1972 Les Houches summer school, only appeared after a prolonged debate (see the reception and publication dates!) [30]. Its results were the following:

- the “black hole uniqueness theorem”, implying the axial symmetry of the configuration and the absence of regular pulsations from black holes,
- the “effective potential technique”, determining the efficiency of the energy emission in the accretion process, and
- the “upper limit on the maximum mass of a neutron star”, discriminating between an unmagnetized neutron star and a black hole.

These three essential components established the paradigm for identifying the first black hole in Cygnus X1. The results were also presented in a widely attended session chaired by John Wheeler at the 1972 Texas Symposium in New York, extensively reported by the *New York Times*. The New York Academy of Sciences, which hosted the symposium, had just awarded me their prestigious Cressy Morrison Award for my work on neutron stars and black holes. Much to their dismay I never handed in the manuscript for the proceedings, since it coincided with a paper submitted for publication [31].

The definition of the paradigm did not come easily but matured slowly after innumerable discussions, mainly on the phone, both with Riccardo Giacconi and Herb Gursky. I still remember an irate professor of the Physics Department at Princeton pointing publicly to my outrageous phone bill of \$274 for one month, at the time considered scandalous, due to my frequent calls to the Smithsonian, and a much more relaxed and sympathetic attitude about this situation by the department chairman Murph Goldberger. This work was finally summarized in two books: one with Herbert Gursky [32], following the 1973 AAAS Annual Meeting in San Francisco, and the second with Riccardo Giacconi [33] following the 1975 LXV Enrico Fermi Summer School (see also the proceedings of the 1973 Solvay Conference).

8 More Analogies Between Thermodynamics and Black Hole Physics

The analogy between thermodynamics and general relativity became in 1971 the topic of the Ph.D. thesis of Wheeler's student Jacob Bekenstein. Through a profound set of gedanken experiments, Jacob pushed further the analogy between thermodynamics and black hole physics. Demetrios and I had formally established the existence of reversible transformations in black hole physics as well as the monotonic increase, as occurs for entropy in thermodynamics, of the irreducible mass M_{irr} of a black hole (from which the word "irreducible" arose). This was formally established independently by Steven Hawking in his area theorem [34]. The complete equivalence between the two results immediately follows from the identity relating the surface area $S = 16\pi M_{\text{irr}}^2$ of the black hole to the irreducible mass M_{irr} , as suggested by Bryce De Witt and confirmed in a quick calculation by Demetrios and myself. Jacob went one step further proposing that the area of the black hole S measured in Planck-Wheeler units could indeed be identified with the entropy [35]. He did this by formulating a statistical interpretation of the black hole entropy, and introducing the first generalized form of the first law of thermodynamics in physical processes involving black holes. Even today, thirty years later, these topics inspire lively debates. Jacob's proposal was extremely interesting and very intriguing at the time and remains so, for me, in some ways even today. It certainly was not contradictory, but I could not find a necessity for transforming it into an identity. The entire matter became the subject of even more intense discussions after Stephen Hawking proposed a physical process which, if true (in the sense used by Wigner), would transform all these theoretical conjectures into physical reality: the black hole quantum evaporation process [36]. Some three decades later, this topic also inspires lively debates. It is likely that these issues will be clarified once there is a theory encompassing both general relativity and relativistic quantum field theories. The basic properties of describing the Hawking radiation process from a black hole can simply be summarized in the following three formulae:

$$\begin{aligned}
 \text{radiation temperature :} & \quad T \simeq 0.62 \times 10^{-7} (M_{\odot}/M) \text{K}, \\
 \text{evaporation time :} & \quad \tau \simeq E/(dE/dt) \sim 2 \times 10^{63} (M/M_{\odot})^3 \text{years}, (4) \\
 \text{energy flux :} & \quad dE/dt \simeq 10^{-22} (M_{\odot}/M)^2 \text{erg/sec}.
 \end{aligned}$$

9 Analogy Between the Electrodynamics of Black Hole and Perfect Conducting Sphere

Before closing this set of analogies I would like to recall a gedanken experiment Richard Hanni and I made by integrating the general relativistic equations back in 1974. The idea was to compute and draw the lines of force of a test charge in the field of a Schwarzschild black hole. This was the topic of the senior thesis of Rick, then my undergraduate student at Princeton. In order to solve this problem we decided to introduce the concept of an “induced charge” on a black hole [37]. By doing so, we proposed the analogy between a black hole and a classical system with a surface equal to the horizon and endowed with an appropriate conductivity (see Figs. 5 and 6). This gedanken experiment opened the way to the field of black hole electrodynamics which was further expanded in important contributions by Damour, Hanni, Ruffini, and Wilson [38] and references therein, and by Thibault Damour. Thibault came to Princeton in 1974 to work with me on his state doctorate thesis to be defended in Paris at the *École Normale Supérieure*. This beautiful thesis incorporated a detailed discussion of the general relativistic effects of electrodynamics of EMBH, including generalized Ohm’s, Joule’s, Ampere’s, Navier Stoke’s laws using as a tool the above-mentioned analogy. Interestingly, this analogy was taken very seriously and written up in the book by Kip Thorne and collaborators [39] as a final theory of black holes (see, however, the subtle relativistic effects and the different conclusions reached by Brian Punsly in his recent interesting book [40]).

The analogy was reconsidered in a widely publicized article on a full page in the science section of the *Frankfurter Allgemeine Zeitung* of January 31, 2001, shown to me by Hagen Kleinert one day after its appearance, praising “new results” in an article by Wilczek and a collaborator [41]. In that paper, Wilczek, without giving any reference, purports to extend our results with Hanni on the induced charge on a Schwarzschild black hole to the case of a Reissner-Nordström geometry with charge Q . They obtained the simple result for the effective potential

$$V(r) = q \left(\frac{Q - q}{r} - \frac{qR}{r^2 - R^2} \right), \quad (5)$$

derived from the technique of introducing the image charge q . Unfortunately, they overlooked that in a very interesting set of papers, Bičák and his students Dvořák [42] and Ledvinka [43] had proven a totally new situation in the lines

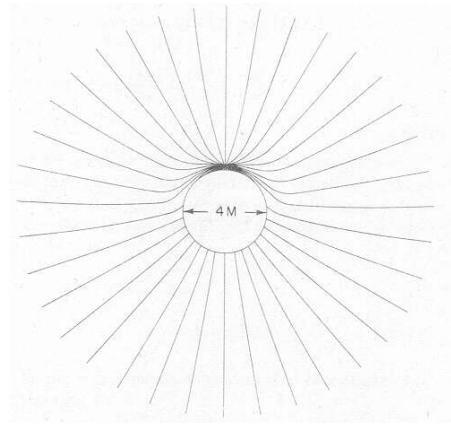


Figure 5. Lines of force of a charge near a black hole (reproduced from Ref. [37]).

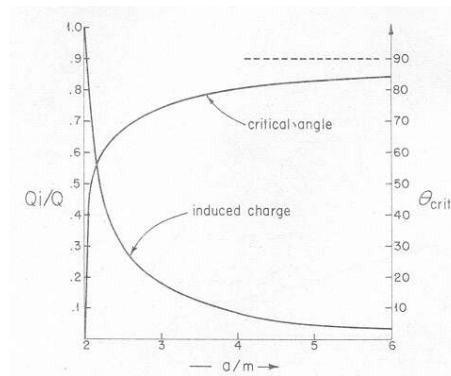


Figure 6. Induced charge on a Schwarzschild black hole (reproduced from Ref. [37]).

of force of a charge in the presence of a Reissner-Nordström geometry (see Fig. 7). In other words, in the limit $Q/M = 1$ of an extreme black hole, no line of force from the test particle crosses the horizon. We are confronted with a totally new effect for the electric field due to general relativity reminiscent of the Meissner effect in classical magnetic fields in the presence of superconductors. Such a general relativistic electric Meissner effect (GREME) does not appear to be contemplated in the Wilczek solution developed in analogy

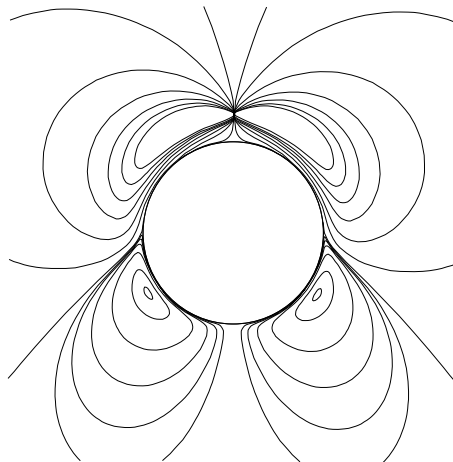


Figure 7. Lines of force of a point charge near an extreme EMBH with $Q = M$ (reproduced from Refs. [42,43]).

to my solution with Hanni. This case can be considered a propedeutic example of the incomparable richness of physical regimes present in the general relativity which, once again, transcends the direct analogies with classical field theories.

This series of events offer a clear pedagogical example of how the enforcement of direct and unproven analogies in general relativity can be dangerous and lead to incorrect conclusions.

10 Discovery of Gamma-Ray Bursts

In 1975 Herbert Gursky and myself had been invited by the AAAS to organize a session on neutron stars, black holes, and binary X-ray sources for the annual meeting in San Francisco. During the preparation of the meeting, we heard that some observations, made from the military Vela satellites and conceived in order to monitor the Limited Test Ban Treaty of 1963 on atomic bomb explosions, had just been unclassified. We asked Ian B. Strong to report, for the first time in a public meeting, on the observation of the gamma-ray bursts [32] (see Fig. 8). From the start it was clear that these signals were not coming either from the earth nor the planetary system. By 1991, a great improvement on the distribution of the gamma-ray bursts came with the

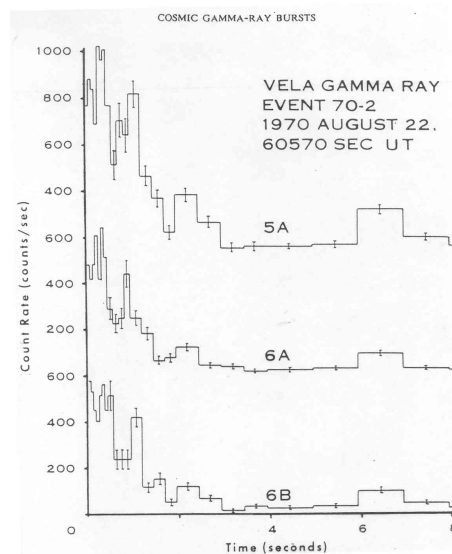


Figure 8. One of the first gamma-ray bursts observed by the Vela satellite (reproduced from an article of Strong in the book [32]).

NASA launch of the Compton Gamma-Ray Observatory which, in ten years of observation, gave a beautiful evidence for the perfect isotropy of the angular distribution of the gamma-ray burst sources in the sky (see Fig. 9). The sources had to be either at cosmological distances or very close to the solar system because one cannot detect the galactic anisotropical distribution there. In the meantime, the number of theories grew exponentially but without any clear conclusion. Quite prominent among these theories were some relating the gamma-ray burst phenomenon to the Hawking radiation process.

11 Analogy of Heisenberg-Euler Critical Capacitor and Vacuum Polarization Around Macroscopic Black Hole

In 1975, following the work of Christodoulou and Ruffini on the energetics of black holes [28], we pointed out [44] the relevance of the vacuum polarization process à la Heisenberg-Euler-Schwinger [45,46] around EMBHs. Such a process can only occur for EMBHs of mass smaller than $7.2 \times 10^6 M_{\odot}$. The basic energetics implications were contained in Table 1 of that paper [44], where it

2704 BATSE Gamma-Ray Bursts

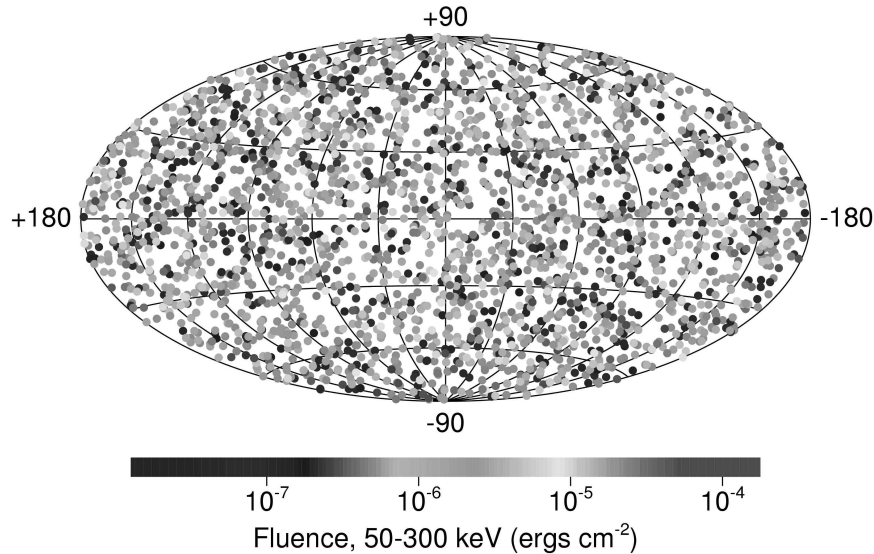


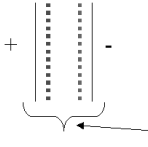
Figure 9. Angular distribution of gamma-ray bursts in galactic coordinates from the Compton GRO satellite.

was also shown that this process is almost reversible in the sense explained by Christodoulou and Ruffini [28], and that it extracts the mass energy of an EMBH very efficiently. We also pointed out that this vacuum polarization process around an EMBH offered a natural mechanism for explaining gamma-ray bursts, in particular their large characteristic energy of 10^{54} ergs (see Fig. 10).

12 Beppo-SAX Satellite and Instantaneous Destruction of More Than 135 Theoretical Models

It was only with the very unexpected and fortuitous observations of the Beppo-SAX satellite that the existence of a long-lasting afterglow of these sources was identified. This led to determining a much more accurate position for these sources in the sky, which permitted for the first time their optical and radio identification. The optical identification has led to the de-

Heisenberg, Euler, 1935 Schwinger, 1951



$$+ \quad - \quad E_c = \frac{m^2 c^3}{\hbar e}; \quad Z_c \sim \frac{\hbar c}{e^2} \sim 137; \quad \Delta t \sim \frac{\hbar}{m_e c^2} \sim 10^{-18} s$$

Damour & Ruffini 1974, "... naturally leads to a most simple model for the explanation of the recently discovered γ -rays bursts"

Kerr-Newmann black hole ($M, Q, a = LM$), with $a^2 + Q^2 \leq M^2$

$$E = Q\Sigma^{-2}(r^2 - a^2 \cos^2 \vartheta) \quad B = Q\Sigma^{-2}2ar \cos \vartheta \quad \Sigma = r^2 + a^2 \cos^2 \vartheta$$

Vacuum polarization process occurs if $3.2M_* \leq M_{BH} \leq 7.2 \cdot 10^6 M_*$.

Maximum energy extractable $1.8 \cdot 10^{54} (M_{BH} M_*)$ ergs

Figure 10. Summary of the EMBH vacuum polarization process (see Ref. [44] for details).

termination of their cosmological distances and to their paramount energetic requirements, in some cases $\geq 10^{54}$ ergs (see Ref. [47]).

The very fortunate interaction and resonance between X-ray, optical, and radio astronomy which, in the seventies, allowed the maturing of the physics and astrophysics of neutron stars and black holes (see e.g. Ref. [33]), promises to be active again in these days in unravelling the physics and astrophysics of the gamma-ray burst sources.

The observations of the Beppo-SAX satellite had a sobering effect on the theoretical development on gamma-ray burst models. Practically all existing theories (see a partial list in Fig. 11) were at once wiped out, not being able to fit the stringent energy requirements imposed by observations. They were particularly constraining for the models based on the Hawking radiation process (see Table 1 for details). These are by far the theoretical predictions furthest from any observational data in the entire history of the homo sapiens and possibly in the entire Universe.

13 Analogy Between Ergosphere and Dyadosphere of Black Holes

The enormous energy requirements of gamma-ray bursts, which are very similar to those predicted by Damour and Ruffini in 1975 [44], have convinced us to return to our earlier work in studying more accurately the process of vac-

Model #	Author	Year	Reference	Main Body	2nd Body	Place	Description
1	Colgate	1966	CJPhys, 46, S476			ST	COS SN shock stellar surface in distant galaxy
2	Colgate	1974	AJ, 187, 333			ST	COS Type II SN shock from, env Comp acc'd at stellar surface
3	Stueckel et al.	1973	Nature, 245, P570			WD	DISK Disk surface from nearby star
4	Stueckel et al.	1973	Nature, 245, P570			WD	DISK Superflare from nearby WD
5	Havel et al.	1973	Nature, 246, P532			WD	DISK Disk surface perturbed to coincide with galactic NS
6	Lamb et al.	1973	Nature, 246, P532			WD	DISK Accretion onto WD from fare in companion
7	Lamb et al.	1973	Nature, 246, P532			WD	DISK Accretion onto NS from fare in companion
8	Lamb et al.	1973	Nature, 246, P532			BH	DISK Accretion onto BH from fare in companion
9	Zuckey	1974	A&S, 107, 111			NS	HALO NS crust cooled by external pressure escapes, explodes
10	Griffith et al.	1974	AJ, 187, 183			DO	SOL Relativistic iron dust grain up-scatters solar radiation
11	Brueker et al.	1974	AJ, 187, 187			DO	SOL "Direct stellar flare on nearby star"
12	Schroeder	1974	SoAstron, 18, 390			WD	COM DISK Comes from system's cloud strikes NS
13	Schroeder	1974	SoAstron, 18, 390			NS	COM Thermal emission when small star heated by SN shock wave
14	Binovitch et al.	1975	ApSS, 35, 23			NS	COS Absorption of neutrino emission from SN in stellar envelope
15	Binovitch et al.	1975	ApSS, 35, 23			SN	COS Thermal emission when small star heated by SN shock wave
17	Paton et al.	1974	Nature, 251, 399			NS	COS Ejected matter from NS explodes
18	Hartik et al.	1974	Nature, 251, 399			NS	COS NS crustal starquake glitch, should time coincide with GRB
19	Channagan	1974	AA, 44, 21			NS	COS White hole emits spectrum that differs with time
20	Priftaki et al.	1975	ApSS, 34, 395			WH	HALO NS corequake causes vibrations, changing E & B fields
21	Priftaki et al.	1975	ApSS, 34, 395			AGN	ST COS Collapses of supermassive body in nucleus of active galaxy
22	Havel et al.	1975	Nature, 256, 112			BH	DISK Infalling accretion disk causes Compton scattering
23	Piran et al.	1975	Nature, 256, 112			BH	DISK Infalling accretion disk causes Compton scattering
24	Channagan	1976	ApSS, 42, 83			WD	DISK Magnetic WD suffers MHD instabilities, fares
25	Mulan	1976	ApSS, 42, 83			WD	DISK Magnetic WD suffers MHD instabilities, fares
26	Woodley et al.	1976	Nature, 263, 101			NS	DISK Carbon dislocation from accreted matter onto NS
27	Lamb et al.	1977	AJ, 214, 268			BH	DISK Inability in accretion onto rapidly rotating BH
28	DeVogela	1978	ApSS, 63, 937			DO	SOL Charged energetic dust grain enters solar system
29	Piran et al.	1980	AA, 87, 224			WD	DISK WD surface nuclear burst causes chromospheric fares
30	Tygan	1980	AA, 87, 224			WD	DISK WD surface nuclear burst causes chromospheric fares
31	Tygan	1980	AA, 87, 224			WD	DISK WD surface nuclear burst causes chromospheric fares
32	Ramaty et al.	1981	ApSS, 75, 193			NS	DISK NS vibrations heat atm to pair produce, annihilate, synch cool
33	Neuman et al.	1980	AJ, 242, 319			NS	AST Asteroid from interstellar medium hits NS
34	Ramaty et al.	1980	Nature, 287, 122			NS	HALO NS core quake caused by phase transition, vibrations
35	Ramaty et al.	1981	AJ, 249, 302			NS	AST Asteroid hits NS, B-field confines mass, creates high temp
36	Mikhailov et al.	1981	ApSS, 77, 469			NS	AST Asteroid hits NS, B-field confines mass, creates high temp
37	Colgate et al.	1981	AJ, 249, 311			NS	AST Asteroid hits NS, B-field confines mass, creates high temp
38	van Buren	1982	CoRP, 20, 72			NS	AST Asteroid enters NS B field, dragged to surface collision
39	Kuznetsov	1982	CoRP, 20, 72			NS	AST Asteroid enters NS B field, dragged to surface collision
40	Kuznetsov	1982	CoRP, 20, 72			NS	AST Asteroid enters NS B field, dragged to surface collision
41	Woodley et al.	1982	AJ, 258, 116			NS	DISK NS fares from pair plasma confined in NS magnetosphere
42	Woodley et al.	1982	AJ, 258, 116			NS	DISK NS fares from pair plasma confined in NS magnetosphere
43	Woodley et al.	1982	AJ, 258, 116			NS	DISK NS fares from pair plasma confined in NS magnetosphere
44	Hameury et al.	1981	AA, 111, 243			NS	DISK Hot fusion runaway on NS B-pole helium lake
45	Mikhailov et al.	1981	AA, 111, 243			NS	DISK Hot fusion runaway on NS B-pole helium lake
46	Hameury et al.	1981	AA, 111, 243			NS	DISK Hot fusion runaway on NS B-pole helium lake
47	Ferraro et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
48	Ferraro et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
49	Ferraro et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
50	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
51	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
52	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
53	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
54	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
55	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
56	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
57	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
58	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
59	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
60	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
61	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
62	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
63	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
64	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
65	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
66	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
67	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
68	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
69	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
70	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
71	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
72	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
73	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
74	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
75	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
76	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
77	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
78	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
79	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
80	Yeh et al.	1982	Nature, 297, 603			NS	DISK E capture trigger B field collapse, B field on NS surface
81	Trofanenko et al.	1989	ApSS, 152, 105			WH	COS Ken-Newman white holes
82	Surroock	1989	AJ, 348, 950			NS	DISK NS E field accelerates electrons which then pair cascade
83	Ferraro et al.	1988	AJ, 335, 171			NS	DISK Narrow absorption features indicate small cold area on NS
84	Rodriguez	1989	AJ, 346, 206			WD	DISK Binary member stars part of orbit, through L1, into primary
85	Pineault et al.	1989	AJ, 347, 1141			NS	COM Disk Fat NS through Dorit clouds, last WD bursts only optical
86	Matta et al.	1989	ApSS, 348, 378			NS	COS Emission of energetic particles from accretion onto NS
87	Trofanenko	1989	ApSS, 159, 301			WH	COS Different types of white, "grey" holes can emit GRB
88	Eichler et al.	1989	ApSS, 159, 301			NS	COS NS - NS binary members collide, coalesce
89	Wang et al.	1989	ApSS, 159, 301			NS	COS NS - NS binary members collide, coalesce
90	Alexander et al.	1989	ApSS, 159, 301			NS	COS NS - NS binary members collide, coalesce
91	Ho et al.	1990	AJ, 351, 601			NS	DISK Cyclic red & blue flash from NS, 40 mV dips, magnetized NS
92	Matta et al.	1990	ApSS, 165, 137			NS	DISK DSD mag resonant cavity in NS atmosphere
93	Darmer	1990	ApSS, 165, 137			COM	DISK NS magnetospheric plasma oscillations
94	Paczynski	1990	ApSS, 165, 137			NS	DISK Beaming of radiation necessary for magnetized neutron stars
95	Paczynski	1990	ApSS, 165, 137			NS	DISK Beaming of radiation necessary for magnetized neutron stars
96	Paczynski	1990	ApSS, 165, 137			NS	DISK Beaming of radiation necessary for magnetized neutron stars
97	Zdziarski et al.	1991	ApSS, 361, 309			NS	DISK Beaming of radiation necessary for magnetized neutron stars
98	Paczynski	1990	ApSS, 165, 137			NS	DISK Beaming of radiation necessary for magnetized neutron stars
99	Trofanenko et al.	1991	ApSS, 361, 309			NS	DISK Beaming of radiation necessary for magnetized neutron stars
100	Matta et al.	1991	ApSS, 361, 309			NS	DISK Beaming of radiation necessary for magnetized neutron stars
101	Hosono et al.	1991	ApSS, 378, 682			NS	DISK NS B field undergoes fast reconnection, accelerates plasma
102	Hameiri et al.	1991	ApSS, 378, 682			NS	DISK NS B field undergoes fast reconnection, accelerates plasma
103	Blais et al.	1991	ApSS, 378, 682			NS	DISK NS B field undergoes fast reconnection, accelerates plasma
104	Franz et al.	1992	ApSS, 385, 145			NS	DISK Strange stars emit binding energy in grav red, and outflow
105	Woodley et al.	1992	ApSS, 385, 145			NS	DISK Strange stars emit binding energy in grav red, and outflow
106	Hogman et al.	1993	ApSS, 411, 541			NS	DISK Low mass X-ray binary evolves into GRB area
107	Dai et al.	1992	ApSS, 386, 164			NS	HALO Accretion WD collapses to NS
108	Thompson et al.	1993	ApSS, 406, 184			NS	HALO NS polar & MW halo boundary specified by hydro density jump
109	Hogman et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
110	Thompson et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
111	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
112	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
113	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
114	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
115	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
116	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
117	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
118	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
119	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
120	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
121	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
122	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
123	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
124	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
125	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
126	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
127	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
128	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
129	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
130	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
131	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
132	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
133	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
134	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump
135	Eichler et al.	1992	ApSS, 386, 164			NS	HALO NS polar & MW halo boundary specified by hydro density jump

Figure 11. Partial list of theories before the Beppo-SAX, from a talk presented at MGIXMM [48].

Table 1. Hawking radiation process versus gamma-ray burst observations. The comparison is performed for a) the energetics $E_{\text{tot}} \simeq 10M_{\odot} \simeq 10^{55}$ erg, b) the time scale $\tau \simeq 1$ sec, and c) the spectrum energy $T = 10^8$ K.

a)

theoretical value	observed value	discrepancy
$T = 6.2 \times 10^{-9}$ K	$T \simeq 10^8$ K	$\sim 10^{-17}$
$\tau = 10^{73}$ sec	$\tau \simeq 1$ sec	$\sim 10^{73}$
$\left(\frac{dE}{dt}\right) = 10^{-24}$ erg/sec	$\left(\frac{dE}{dt}\right) \simeq 10^{54}$ erg/sec	$\sim 10^{-78}$

b)

theoretical value	observed value	discrepancy
$E_{\text{tot}} \simeq 10^{-24} M_{\odot} \simeq 10^{30}$ erg	$E_{\text{tot}} \simeq 10^{55}$ erg	$\sim 10^{-25}$
$T = 10^{17}$ K	$T \simeq 10^8$ K	$\sim 10^9$
$\left(\frac{dE}{dt}\right) = 10^{26}$ erg/sec	$\left(\frac{dE}{dt}\right) \simeq 10^{54}$ erg/sec	$\sim 10^{-28}$

c)

theoretical value	observed value	discrepancy
$E_{\text{tot}} \simeq 10^{-9} M_{\odot} \simeq 10^{45}$ erg	$E_{\text{tot}} \simeq 10^{55}$ erg	$\sim 10^{-10}$
$\tau = 10^{43}$ sec	$\tau \simeq 1$ sec	$\sim 10^{43}$
$\left(\frac{dE}{dt}\right) = 10^{-4}$ erg/sec	$\left(\frac{dE}{dt}\right) \simeq 10^{54}$ erg/sec	$\sim 10^{-58}$

uum polarization and the region of pair creation around an EMBH. This has led a) to the new concept of the dyadosphere of an EMBH (“dyado” from the Greek word for “pair”), b) to the concept of a plasma-electromagnetic (PEM) pulse, and c) to the analysis of its temporal evolution generating signals with the characteristic features of a gamma-ray burst.

In our theoretical approach, we claim that, by the observations of gamma-ray bursts, we are witnessing the formation of an EMBH, and that we thus follow the process of a gravitational collapse in real time. Even more important: the tremendous energies involved in the energetics of these processes

coincide with the extractable energy of black holes which is described by the Eqs. (1)–(3).

Various models have been proposed in order to extract the rotational energy of black holes by processes of relativistic magnetohydrodynamics (see e.g. Ref. [49]). It should be expected, however, that these processes are relevant over the long time scales characteristic of accretion processes.

In the present case of gamma-ray bursts, a sudden mechanism appears to be at work on time scales of the order of few seconds or shorter, and they are naturally explained by the vacuum polarization process introduced by Damour and Ruffini [44].

The fundamental new points we have found reexamining our previous work can be summarized as follows (see Ref. [50] for details):

- The vacuum polarization process can occur in an extended region around the black hole called the dyadosphere, extending from the horizon radius r_+ out to the dyadosphere radius r_{ds} . Only black holes with a mass larger than the upper limit of a neutron star and up to a maximum mass of $7.2 \times 10^6 M_\odot$ can have a dyadosphere.
- The efficiency of transforming the mass-energy of a black hole into particle-antiparticle pairs outside the horizon can approach 100% for black holes in the above mass range.
- The pairs created are mainly positron-electron pairs and their number is much larger than the quantity Q/e one would have naively expected on the grounds of qualitative considerations. It is actually given by $N_{\text{pairs}} \sim Q/e \times r_{ds}/(\hbar/mc)$, where m and e are the electron mass and charge, respectively. The energy of the pairs, and consequently the emission of the associated electromagnetic radiation, peaks in the gamma-X-ray region, as a function of the black hole mass.

Let us now recall the main results on the dyadosphere obtained by Preparata, Ruffini, and Xue [50]. Although the general considerations presented by Damour and Ruffini [44] did refer to a Kerr-Newmann field with axial symmetry around the rotation axis, we have there considered, for simplicity, the case of a non-rotating Reissner-Nordström EMBH to illustrate the basic gravitational-electrodynamical process. The dyadosphere then lies between the radius

$$r_{ds} = \left(\frac{\hbar}{mc}\right)^{\frac{1}{2}} \left(\frac{GM}{c^2}\right)^{\frac{1}{2}} \left(\frac{m_p}{m}\right)^{\frac{1}{2}} \left(\frac{e}{q_p}\right)^{\frac{1}{2}} \left(\frac{Q}{\sqrt{GM}}\right)^{\frac{1}{2}} \quad (6)$$

and the horizon radius

$$r_+ = \frac{GM}{c^2} \left(1 + \sqrt{1 - \frac{Q^2}{GM^2}} \right). \quad (7)$$

The number density of pairs created in the dyadosphere is

$$N_{e^+e^-} \simeq \frac{Q - Q_c}{e} \left[1 + \frac{(r_{ds} - r_+)}{\hbar/mc} \right], \quad (8)$$

where $Q_c = 4\pi r_+^2 m^2 c^3 / \hbar e$. The total energy of pairs converted from the static electric energy and deposited within the dyadosphere is then

$$E_{e^+e^-}^{\text{tot}} = \frac{1}{2} \frac{Q^2}{r_+} \left(1 - \frac{r_+}{r_{ds}} \right) \left[1 - \left(\frac{r_+}{r_{ds}} \right)^2 \right]. \quad (9)$$

There are many interesting analogies between the ergosphere and the dyadosphere:

- Both of them are extended regions around the black hole.
- In both regions the energy of the black hole can be extracted, approaching the limiting case of reversibility from Christodoulou and Ruffini [28].
- The electromagnetic energy extraction by the pair creation process in the dyadosphere is much simpler and less contrived than the corresponding process of extracting rotational energy from the ergosphere.

14 Analogies Between EM Pulse of Atomic Explosion and PEM Pulse of Black Holes

The analysis of the radially resolved evolution of the energy deposited within the e^+e^- -pair and photon-plasma fluid created in the dyadosphere of an EMBH is much more complex than we had initially anticipated. In this respect, the collaboration with Jim Wilson and his group at Livermore Radiation Laboratory was very important for us. Recently, we decided to join forces in a new collaboration, renewing our previous successful collaboration of 1974 [49]. We proceeded in parallel: in Rome with simple, almost analytic models, in Livermore by confirming these models with computer codes [51,52].

For the evolution we assumed the relativistic hydrodynamic equations (for details see Refs. [53,54]). We assumed the plasma fluid of e^+e^- -pairs,

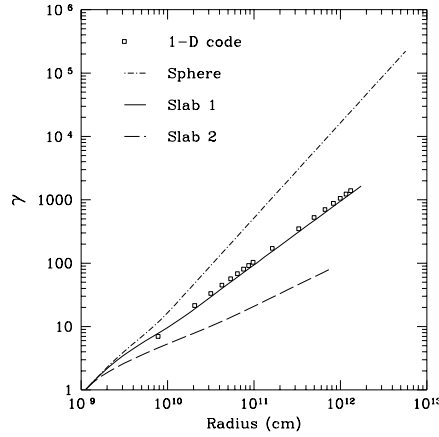


Figure 12. Lorentz factor γ as a function of radius. Three models for the expansion pattern of the e^+e^- pair plasma are compared with the results of the one-dimensional hydrodynamic code for a $1000M_\odot$ black hole with charge $Q = 0.1Q_{\max}$. The 1-D code has an expansion pattern that strongly resembles that of a shell with constant coordinate thickness (reproduced from Ref. [54]).

photons and baryons to be a simple perfect fluid in curved space-time. The baryon-number and energy-momentum conservation laws are

$$(n_B U^\mu)_{;\mu} = (n_B U^t)_{;t} + \frac{1}{r^2} (r^2 n_B U^r)_{;r} = 0, \quad (10)$$

$$(T_\mu^\sigma)_{;\sigma} = 0, \quad (11)$$

with the rate equation

$$(n_{e^\pm} U^\mu)_{;\mu} = \bar{\sigma} v [n_{e^-}(T) n_{e^+}(T) - n_{e^-} n_{e^+}], \quad (12)$$

where U^μ is the four-velocity of the plasma fluid, n_B is the proper baryon-number density, n_{e^\pm} are the proper densities of electrons and positrons e^\pm , σ is the mean pair annihilation-creation cross-section, v is the thermal velocity of e^\pm , and $n_{e^\pm}(T)$ are the proper number-densities of e^\pm at an appropriate equilibrium temperature T . The calculations are continued until the plasma fluid expands, cools, and the e^+e^- pairs recombine, making the system optically thin.

The results of the Livermore computer work are compared and contrasted with three almost analytical models:

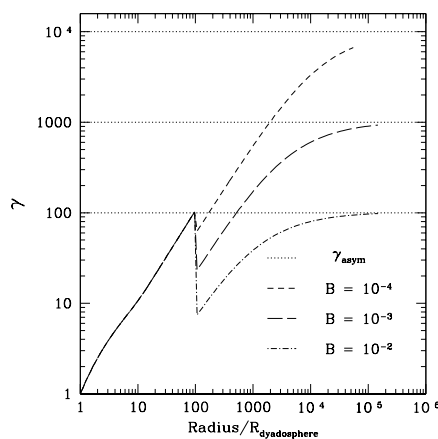


Figure 13. Lorentz factor γ as a function of radius for the PEM pulse interacting with the baryonic matter of the remnant (PEMB pulse) for selected values of the baryonic matter (reproduced from Ref. [55]).

- spherical model where the radial component of four-velocity is of the form $U(r) = Ur/\mathcal{R}$ with U being the four-velocity at the surface \mathcal{R} of the plasma, similar to a portion of a Friedmann model,
- slab 1 where $U(r) = U_r = \text{const.}$, which is an expanding slab with constant width $\mathcal{D} = R_o$ in the coordinate frame in which the plasma is moving,
- slab 2 which is an expanding slab with constant width $R_2 - R_1 = R_o$ in the comoving frame of the plasma.

We compute the relativistic Lorentz factor γ of the expanding e^+e^- pair and photon plasma. In Fig. 12 we see a comparison of the Lorentz factor of the expanding fluid as a function of radius for all these models. We can see that the one-dimensional computation, of which only a few significant points are plotted, matches the expansion pattern of a shell of constant coordinate thickness.

In analogy with the notorious electromagnetic radiation EM pulse of some explosive events, we called this relativistic counterpart of an expanding pair electromagnetic radiation shell a PEM pulse. In recent works we have computed the interaction of the expanding plasma with the surrounding baryonic matter [55] (see Fig. 13). We have also been able to follow the expansion

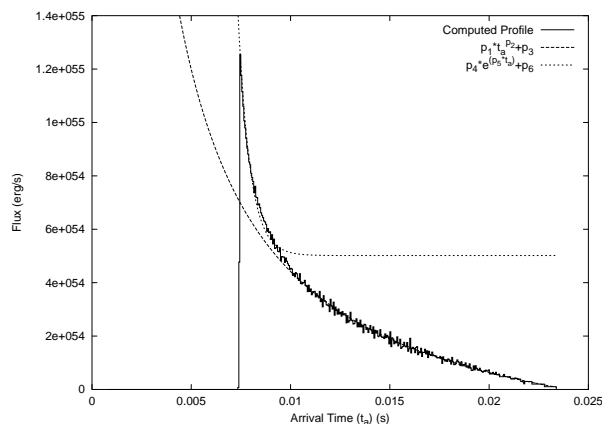


Figure 14. Proper gamma-ray burst from an EMBH with $M = 100M_{\odot}$ and $Q = 0.1Q_{\max}$ (reproduced from Ref. [56]).

process all the way to the point where the transparency condition is reached and what we have defined the “proper gamma-ray burst” is emitted [56] (see Fig. 14). These results of our theoretical model have reached the point where they can be subjected to a direct comparison with observational data.

15 Three New Paradigms for Interpreting Gamma-RAY Bursts

Starting from this theoretical background, we have moved ahead to fit, for the first time, the observational data on the ground of the EMBH model. We have used, as a prototype, the object GRB 991216, both for its very high energetics, which we have estimated in the range of $E_{\text{dya}} \sim 9.57 \times 10^{52}$ ergs, and for the superb data obtained by the Chandra and RXTE satellites. We have found a necessity to formulate in our novel approach three new paradigms, in order to understand the gamma-ray burst phenomenon:

- the relative space-time transformations (RSTT) paradigm (see Ref. [57]),
- the interpretation of the burst structure (IBS) paradigm (see Ref. [58]),
- the multiple-collapse time sequence (MCTS) paradigm (see Ref. [59]).

These results are currently under refereeing process in *Astrophysical Journal Letters* since 28/11/2000.

16 Conclusions

From the above experience, I can venture to formulate some conclusions, which may be of general validity.

16.1 *On Analogies*

- Analogies have been extremely helpful in establishing similarities and deepening our physical understanding, if applied to two circumstances both derived within a general relativistic framework. The analogies between dyadosphere and ergosphere are good examples.
- Analogies between classical regimes and general relativistic regimes have sometimes been helpful in giving a glance at the enormous richness of new physical processes contained in Einstein's theory of space-time structure. In some cases they have allowed us to reach new knowledge and formalize new physical laws, the derivation of Eqs. (1)–(3) being a good example. Such analogies have also dramatically evidenced the enormous differences in depth and physical complexity between classical physics and general relativistic effects. The extraction of rotational energy from a neutron star and a rotating black hole are good examples.
- In no way an analogy based on classical physics can be enforced for general relativistic regimes. Such an analogy is too constraining, and the relativistic theory shows systematically a wealth of novel physical circumstances and conceptual subtleties which are unreachable within a classical theory. The analogies in classical electrodynamics we just outlined are good examples.

16.2 *On New Paradigms*

The establishment of new paradigms is essential to the scientific process, and certainly not easy to obtain. Such paradigms are important in order to guide a meaningful comparison between theories and observations, and much attention should be given to their development and inner conceptual consistency.

16.3 *On Observational Data*

The major factors in driving the progress of scientific knowledge are always the confrontation of theoretical predictions with observational data. In recent

years, the evolution of new technologies has allowed to dramatically improve the sensitivity of observational apparatus. It is gratifying that in this process of learning the structure of our Universe, observational data do not intervene in a marginal way, but with clear and unequivocal results: they confirm by impressive agreement the correct theories and disprove, by equally impressive disagreement, the wrong ones. Our examples are also significant features of this process.

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References

- [1] R.P. Feynman and H. Kleinert, *Phys. Rev. A* **34**, 5080 (1986).
- [2] H. Kleinert, *Gauge Fields in Condensed Matter*, Vol. I: *Superflow and Vortex Lines* and Vol. II: *Stresses and Defects* (World Scientific, Singapore, 1989).
- [3] H. Kleinert, *Path Integrals in Quantum Mechanics, Statistics, and Polymer Physics*, 2nd ed. (World Scientific, Singapore, 1995).
- [4] H. Kleinert and V. Schulte-Frohlinde, *Critical Properties of ϕ^4 -Theories* (World Scientific, Singapore, 2001).
- [5] E. Fermi, *Rend. Acc. Lincei* **6**, 602 (1928).
- [6] L.D. Landau, *Phys. Zeitschr. Sov.* **1**, 285 (1932).
- [7] S. Chandrasekhar, *Introduction to the Theory of the Stellar Structure* (University of Chicago Press, Chicago, 1931).
- [8] J.R. Oppenheimer and G.M. Volkoff, *Phys. Rev.* **55**, 374 (1939).
- [9] J.R. Oppenheimer and H. Snyder, *Phys. Rev.* **56**, 455 (1939).
- [10] C. Størmer, *Astrophysica Norvegica* **1**, 1 (1934).
- [11] B. Carter, *Phys. Rev.* **174**, 1559 (1968).
- [12] M. Rees, R. Ruffini, and J.A. Wheeler, *Black Holes, Gravitational Waves, and Cosmology* (Gordon and Breach, New York, 1974); also in Russian (MIR, Moscow, 1973).
- [13] R. Ruffini and J.A. Wheeler, *Physics Today* **24**, 30 (1971).
- [14] T. Regge and J.A. Wheeler, *Phys. Rev.* **108**, 1063 (1957).
- [15] F.J. Zerilli, *Phys. Rev. D* **2**, 2141 (1970).
- [16] F.J. Zerilli, *Phys. Rev. D* **9**, 860 (1974).

- [17] S.A. Teukolsky, *Astrophys. J.* **185**, 635 (1973).
- [18] C.H. Lee, *J. Math. Phys.* **17**, 1226 (1976); *Prog. Theor. Phys.* **66**, 180 (1981).
- [19] S. Chandrasekhar, *The Mathematical Theory of Black Holes* (Clarendon Press, Oxford, 1983); *Proc. R. Soc. Lond. A* **349**, 571 (1976).
- [20] G.L. Cruciani, in *Proceedings of the Third Iera Network Workshop*, Eds. C. Cherubini and R. Ruffini, *Nuovo Cim. B* **115**, 693 (2000).
- [21] C. Cherubini and R. Ruffini, in *Proceedings of the Third Iera Network Workshop*, Eds. C. Cherubini and R. Ruffini, *Nuovo Cim. B* **115**, 699 (2000).
- [22] D. Bini, C. Cherubini, R.T. Jantzen, and R. Ruffini, *Wave Equation for Tensor Valued p-Forms: Application to the Teukolsky Master Equation*, preprint 2001.
- [23] D. Bini, C. Cherubini, R.T. Jantzen, and R. Ruffini, *De Rham Wave Equation for the Gravitational and Electromagnetic Fields in Vacuum*, preprint 2001.
- [24] R. Penrose, *Nuovo Cim. Rivista* **1**, 252 (1969).
- [25] R. Ruffini and J.A. Wheeler, in *Proceedings of the Conference on Space Physics*, Eds. V. Hardy and H. Moore (E.S.R.O., Paris, 1971). The preparation of this report took more than one year and the authors were unwilling to publish parts of it before the final publication. In order to avoid delays, the results of the energy extraction process from a Kerr black hole, as well as the definition of the “ergosphere”, were inserted as Fig. 2 in the article of Christodoulou [27] which was published on November 30, 1970.
- [26] R.M. Floyd and R. Penrose, *Nature* **229**, 177 (1971), submitted December 16, 1970.
- [27] D. Christodoulou, *Phys. Rev. Lett.* **25**, 1596 (1970).
- [28] D. Christodoulou and R. Ruffini, *Phys. Rev. D* **4**, 3552 (1971).
- [29] B. de Witt and C. de Witt (Eds.), *Black Holes* (Gordon and Breach, New York, 1973).
- [30] C. Rhoades and R. Ruffini *Phys. Rev. Lett.* **32**, 324 (1974).
- [31] R.W. Leach and R. Ruffini, *Astrophys. J.* **180**, L15 (1973).
- [32] H. Gursky and R. Ruffini, Eds. and coauthors, *Neutron Stars, Black Holes and Binary X-Ray Sources* (D. Reidel, Dordrecht, 1975).
- [33] R. Giacconi and R. Ruffini, Eds. and coauthors, *Physics and Astrophysics of Neutron Stars and Black Holes* (North Holland, Amsterdam, 1978).

-
- [34] S.W. Hawking, *Phys. Rev. Lett.* **26**, 1344 (1971).
 - [35] J.D. Bekenstein, *Phys. Rev. D* **7**, 2333 (1973).
 - [36] S.W. Hawking, *Nature* **248**, 30 (1974).
 - [37] R.S. Hanni and R. Ruffini, *Phys. Rev. D* **8**, 3259 (1973).
 - [38] T. Damour, R.S. Hanni, R. Ruffini, and J. Wilson, *Phys. Rev. D* **17**, 1518 (1978).
 - [39] K.S. Thorne, R.H. Price, and D.A. Macdonald, *Black Holes: The Membrane Paradigm* (Yale University Press, New Haven, 1986).
 - [40] B. Punsly, *Black Hole GravitoHydro-Magnetics* (Springer, Berlin, 2001).
 - [41] F. Wilczek and M.K. Parikh, *Phys. Rev. Lett.* **85**, 5042 (2000).
 - [42] J. Bičák and L. Dvořák, *Phys. Rev. D* **22**, 2933 (1980).
 - [43] J. Bičák and T. Ledvinka, in *Proceedings of the Third Iera Network Workshop*, Eds. C. Cherubini and R. Ruffini, *Nuovo Cim. B* **115**, 739 (2000).
 - [44] T. Damour and R. Ruffini, *Phys. Rev. Lett.* **35**, 463 (1975).
 - [45] W. Heisenberg and H. Euler, *Z. Phys.* **98**, 714 (1935).
 - [46] J. Schwinger, *Phys. Rev.* **98**, 714 (1951).
 - [47] E. Costa, invited talk at *IX Marcel Grossmann Meeting on General Relativity*, Eds. V. Gurzadyan, R.T. Jantzen, and R. Ruffini (World Scientific, Singapore, 2001).
 - [48] V. Gurzadyan, R.T. Jantzen, and R. Ruffini (Eds.), *Proceedings of IX Marcel Grossmann Meeting on General Relativity* (World Scientific, Singapore, 2001).
 - [49] R. Ruffini and J.R. Wilson, *Phys. Rev. D* **12**, 2959 (1975).
 - [50] G. Preparata, R. Ruffini, and S.-S. Xue, preprint (1998); *Astronomy and Astrophysics* **338**, L87 (1998).
 - [51] J.R. Wilson, J.D. Salmonson, and G.J. Mathews, in *Gamma-Ray Bursts: 4th Huntsville Symposium*, Eds. C.A. Meegan, R.D. Preece, and T.M. Koshut (A.I.P., New York, 1998).
 - [52] J.R. Wilson, J.D. Salmonson, and G.J. Mathews, in *2nd Oak Ridge Symposium on Atomic and Nuclear Astrophysics*, Ed. A. Mezzacappa (IOP Publishing, London, 1998).
 - [53] R. Ruffini, J.D. Salmonson, J.R. Wilson, and S.-S. Xue, *Astronomy and Astrophysics Suppl. Ser.* **138**, 511 (1998).
 - [54] R. Ruffini, J.D. Salmonson, J.R. Wilson, and S.-S. Xue, *Astronomy and Astrophysics* **350**, 334 (1999).
 - [55] R. Ruffini, J.D. Salmonson, J.R. Wilson, and S.-S. Xue (2000), *Astronomy and Astrophysics* **359**, 855 (2000).

- [56] C.L. Bianco, R. Ruffini, and S.-S. Xue, *Astronomy and Astrophysics* **368**, 377 (2001).
- [57] R. Ruffini, C.L. Bianco, P. Chardonnet, F. Fraschetti, and S.-S. Xue, submitted to *Astrophys. J. Lett.*
- [58] R. Ruffini, C.L. Bianco, P. Chardonnet, F. Fraschetti, and S.-S. Xue, submitted to *Astrophys. J. Lett.*
- [59] R. Ruffini, C.L. Bianco, P. Chardonnet, F. Fraschetti, and S.-S. Xue submitted to *Astrophys. J. Lett.*